

13장 만유인력

4. $m_1 + m_2 = 5.00 \text{ kg}$ 이므로, $m_2 = 5.00 \text{ kg} - m_1$ 이용

$$F = G \frac{m_1 m_2}{r^2} = 1.00 \times 10^{-8} \text{ N}$$

$$1.00 \times 10^{-8} \text{ N} = (6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2) \frac{m_1 (5.00 \text{ kg} - m_1)}{(0.200 \text{ m})^2}$$

$$(5.00 \text{ kg})m_1 - m_1^2 = \frac{(1.00 \times 10^{-8} \text{ N})(0.0400 \text{ m}^2)}{6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2} = 6.00 \text{ kg}$$

$$m_1^2 - (5.00 \text{ kg})m_1 + 6.00 \text{ kg} = 0, \quad (m_1 - 3.00 \text{ kg})(m_1 - 2.00 \text{ kg}) = 0$$

$$m_1 = 3.00 \text{ kg}, m_2 = 2.00 \text{ kg} \quad \text{또는} \quad m_1 = 2.00 \text{ kg}, m_2 = 3.00 \text{ kg}$$

6. $g_c = G \frac{M_e}{R_e^2}$, $g_m = G \frac{M_m}{R_m^2}$, $g_m = \frac{1}{6} g_e$, $R_m = 0.250 R_e$

$$\frac{M_m}{(0.250 R_e)^2} = \frac{1}{6} \frac{M_e}{R_e^2} \quad \therefore M_m = \frac{1}{6} (0.250)^2 M_e$$

$$\frac{\rho_m}{\rho_e} = \frac{M_m/R_m^3}{M_e/R_e^3} = \frac{\frac{1}{6} (0.250)^2 M_e / (0.250 R_e)^3}{M_e/R_e^3} = \frac{(0.250)^2}{6(0.250)^3} = \frac{1}{6(0.250)} = \frac{2.00}{3} = 0.667$$

9. $F_g = mg = G \frac{M_E m}{r^2}$ 이용,

$$g = G \frac{M_E}{r^2} \text{ 에서 중력 가속도가 } 9.00 \text{ m/s}^2 \text{ 이므로,}$$

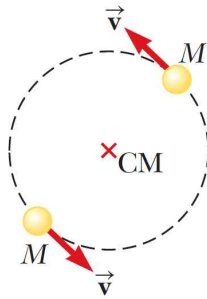
$$r = \sqrt{\frac{GM_E}{g}} = \sqrt{\frac{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2)(5.98 \times 10^{24} \text{ kg})}{9.00 \text{ m/s}^2}} = 6.66 \times 10^6 \text{ m}$$

케플러의 제 3법칙을 이용하여,

$$\begin{aligned} T &= 2\pi \sqrt{\frac{r^3}{GM_E}} = 2\pi \sqrt{\frac{(6.66 \times 10^6 \text{ m})^3}{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2)(5.98 \times 10^{24} \text{ kg})}} \\ &= 5.41 \times 10^3 \text{ s} \end{aligned}$$

$$T = (5.41 \times 10^3 \text{ s}) \left(\frac{1 \text{ h}}{3600 \text{ s}} \right) = 1.50 \text{ h} = 90.0 \text{ min}$$

11.



별간의 만유인력이 구심력이 되므로

$$f = M \frac{v^2}{r} = G \frac{M^2}{(2r)^2} \dots \textcircled{1}$$

공전주기 $T = 14.4 \times 24 \times (60)^2$ s 이고,

$$vT = 2\pi r \text{ 이므로, } r = \frac{vT}{2\pi} \dots \textcircled{2}$$

①과 ②로부터, $M = \frac{4rv^2}{G} = \frac{2v^3 T}{\pi G} = 1.26 \times 10^{32}$ kg

17. 초기속력 $v_i = 2.00 \times 10^4$ m/s , 만유인력 상수 $G = 6.674 \times 10^{-11}$ N • m²/kg²

지구질량 $M_e = 5.97 \times 10^{24}$ kg , 지구반지름 $R_e = 6.37 \times 10^6$ m

지구-로켓 계에 대해 $E_i = E_f$ 로부터,

$$\frac{1}{2}mv_i^2 - \frac{GM_e m}{R_e} = \frac{1}{2}mv_f^2 \quad \therefore v_f = \sqrt{v_i^2 - \frac{2GM_e}{R_e}} = 1.66 \times 10^4 \text{ m/s}$$